



# Future neutrino beams(>2015): $\beta$ -beams and Neutrino Factories

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Pasquale Migliozzi

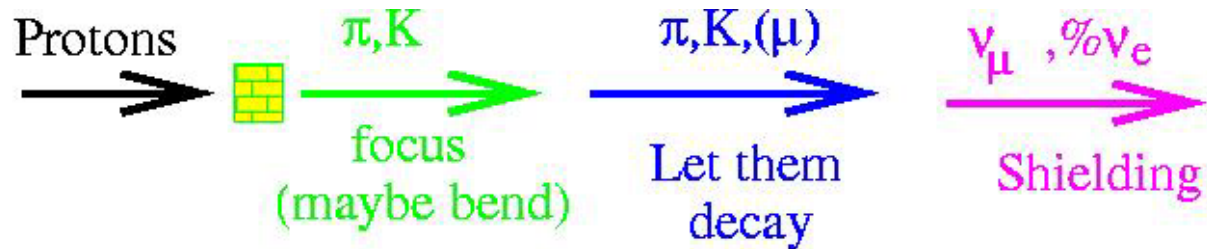


Why conventional  $\nu$  beams are not  
adequate to precision measurements  
of the PMNS matrix?



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# Working principle of a conventional neutrino beam



High energy protons ( $O(10 \text{ GeV})$  to  $O(100 \text{ GeV})$ ) are sent onto a target (typically Be, graphite) where  $\pi$  and K are produced copiously

$\pi^+/K^+$  ( $\pi^-/K^-$ ) are focused and  $\pi^-/K^-$  ( $\pi^+/K^+$ ) are defocused in order to obtain a  $\nu_\mu$  (anti- $\nu_\mu$ ) beam through a magnetic lens system

Focused hadrons are let decay into a "decay tunnel" to produce neutrinos of the wanted flavor (and not only)

A shielding is put at the end of the decay tunnel in order to absorb charged particles associated with the neutrino beam (this step is only true for expts located very close ( $L < 1 \text{ km}$ ) to the neutrino source)



## Problems with conventional neutrino beams in predicting fluxes and composition (I)

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- Description of the proton beam
- Particle yield in the p-Target interaction
- Description of the focusing system
- Description of the particle trajectories after the focusing system (important to extrapolate from the decay point to the detector location)



## Problems with conventional neutrino beams in predicting fluxes and composition (II)

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Use standard MC simulation : Geant, Fluka, Mars to full simulate target production + beamline (SLOW)

Use dedicated parametrizations for secondary production in target ( Sanford-Wang, Malensek, BMPT .... based on available data (Na56/SPY, ...)) + simulation of beamline (FAST)

Secondary components -> needs better knowledge of hadron production in the target

More data needed on hadroproduction

Two possible solutions

- NOMAD approach (no close detector)
- One (or even more) detector(s) approach (near to far extrapolation)



# Systematic uncertainties

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- Uncertainty on the yields of particles from p-Be interactions
- Uncertainty on the yields of particles interactions other than p-Be
- Proton interactions downstream of the Be target
- Reinteractions in the Be target and downstream of the target
- Position and angular divergence of p beam
- Magnetic field in the horn and the reflector
- Inaccuracies in the simulation of the beam line elements
- Misalignment of the beam line elements



# The BETA-BEAM

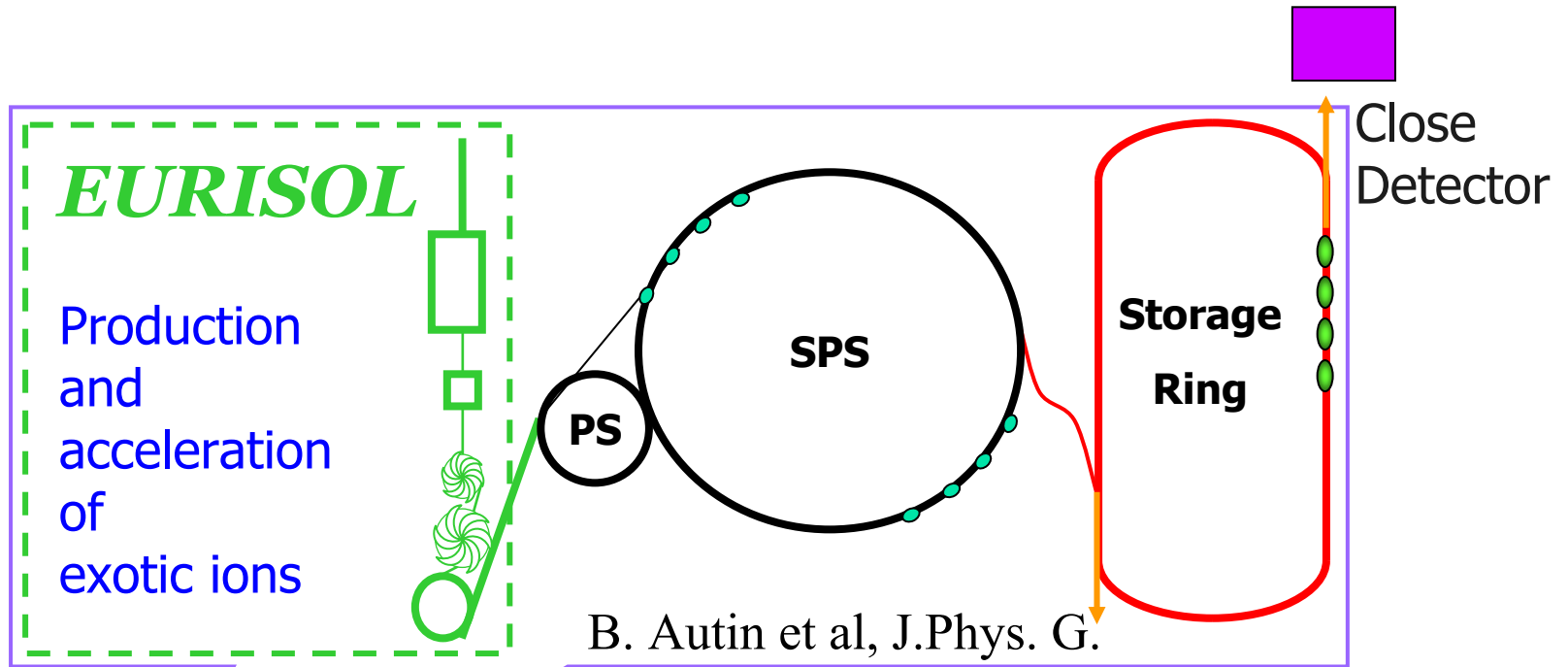
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The BB was born in 2001 when P. Zucchelli put forward the idea to produce pure (anti-) $\nu_e$  beam from the decay of radioactive ions

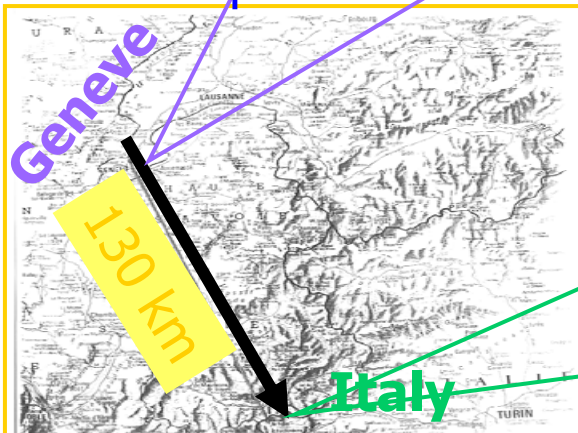
1. Produce a radioactive ion with a short beta-decay lifetime
2. Accelerate the ion in a conventional way (PS) to "high" energy
3. Store the ion in a decay ring with straight sections
4. By its  $\beta$ -decay,  $\nu_e$  (anti- $\nu_e$ ) will be produced

- SINGLE flavour ( $\nu_e$ )
- Known spectrum/intensity
- Focussed ( $1/\gamma$ )
- The energy depends on the ion  $\gamma$

# Baseline Scenario at CERN



One of the possible scenarios :



**UNO**



CP VIOLATION,  
PROTON DECAY  
SUPERNOVA v.

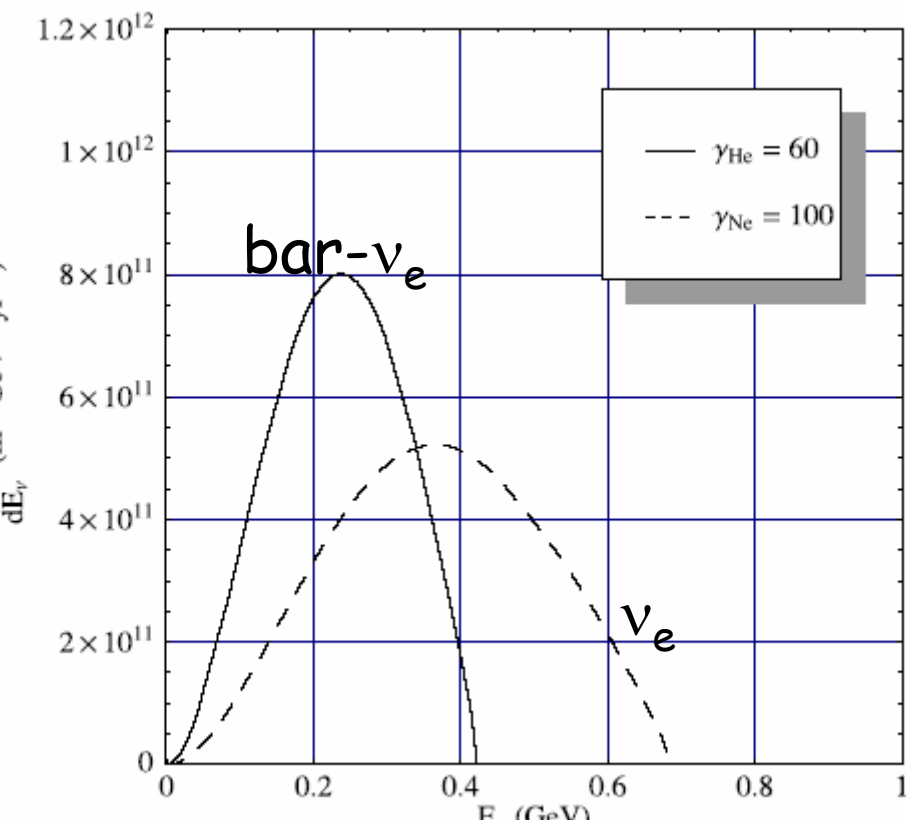


# BB fluxes at Fréjus location

$$\gamma(^6\text{He}) = 60 \quad \gamma(^{18}\text{Ne}) = 100$$

$$\langle E \rangle = 230 \text{ MeV} \quad \langle E \rangle = 370 \text{ MeV}$$

All  $^{18}\text{Ne}$  decay modes are accounted for



Element	End-Point (MeV)	Decay Fraction
$^{18}\text{Ne}$	34.114	92.1%
	23.699	7.7%
	17.106	0.2%
$^6\text{He}$	35.078	100%

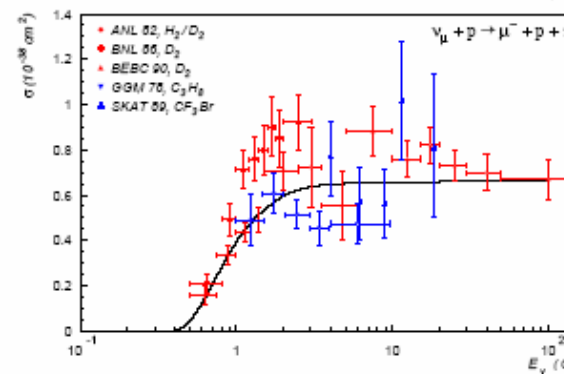
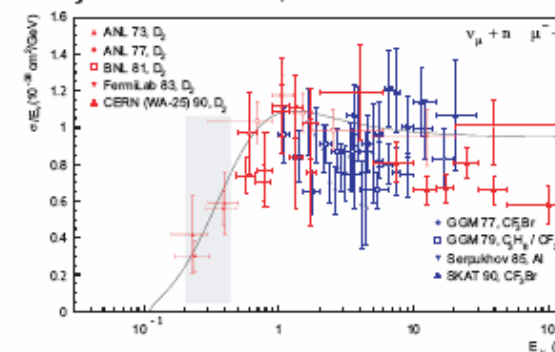
A  $m_e \neq 0$  is considered, unlike in hep-ex/0310059. Although it seems negligible, this has a strong impact on the expected rate due to the strong cross-section suppression at small energies.

Rate( $m_e = 0$ ) smaller than 20% wrt Rate( $m_e \neq 0$ )!

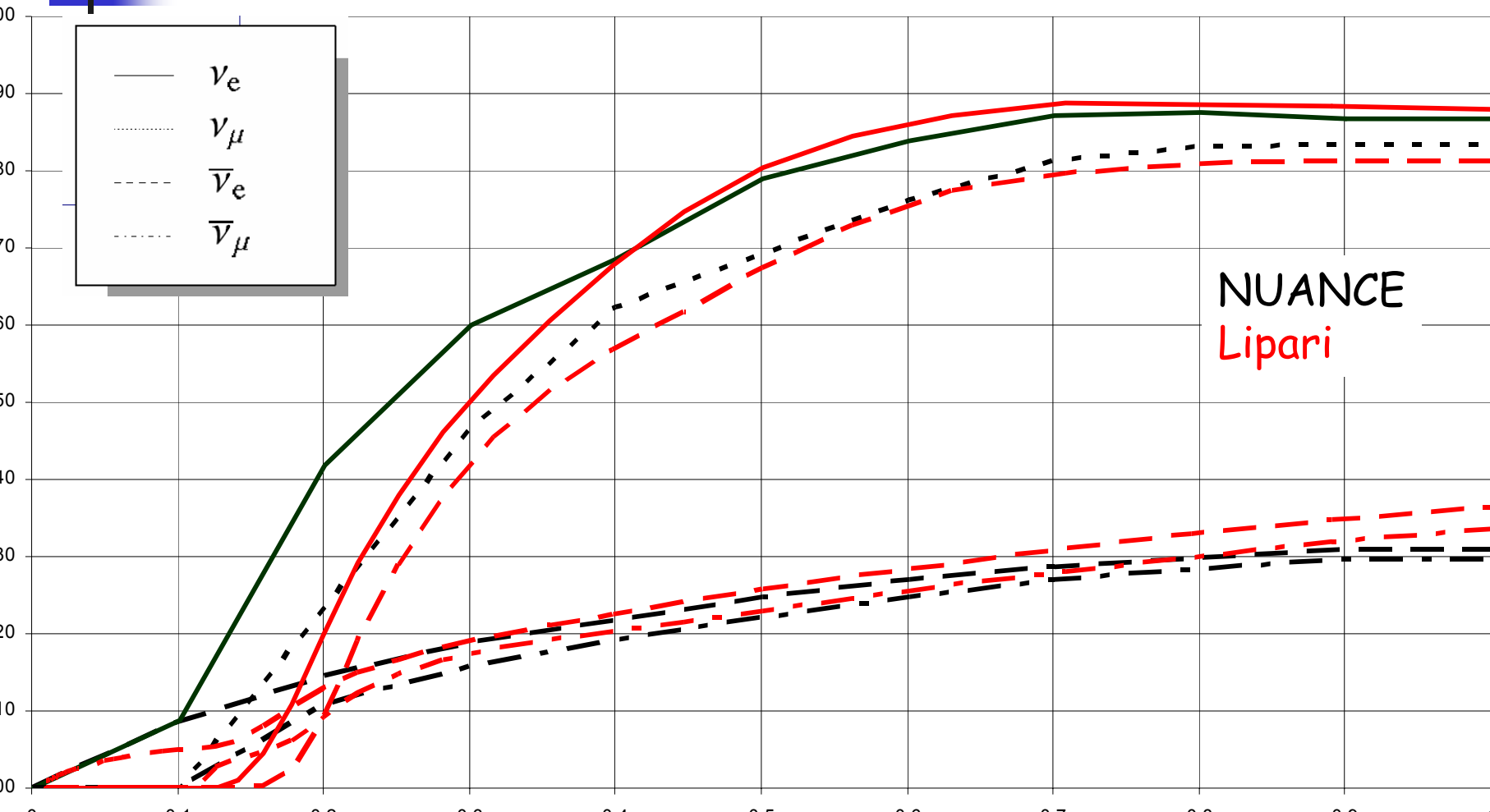
# The cross-section problem

- The present knowledge of neutrino and anti-neutrino cross-sections is rather poor below 1 GeV (see plot)
- On top of that, the few available data are not on water
  - Very difficult the extrapolation from different nuclei due to nuclear effects
- Therefore, it is not astonishing that different calculations can differ up to a factor 2
- In the following we compare two calculations on water: NUANCE and one from P. Lipari (adopted in this work)

V.V. Lyubushkin et al., internal NOMAD memo



# Cross-section comparison





# Target values for the decay ring

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## ${}^6\text{Helium}^{2+}$

- In Decay ring:  $1.0 \times 10^{14}$  ions
- Energy: 139 GeV/u
- Rel. gamma: 150
- Rigidity: 1500 Tm

## ${}^{18}\text{Neon}^{10+}$ (single target)

- In decay ring:  $4.5 \times 10^{12}$  ions
- Energy: 55 GeV/u
- Rel. gamma: 60
- Rigidity: 335 Tm

- The neutrino beam at the experiment should have the “time stamp” of the circulating beam in the decay ring.
- The beam has to be concentrated to as few and as short bunches as possible to maximize the number of ions/nanosecond. (background suppression), aim for a duty factor of  $10^{-4}$



# Why high $\gamma$ BB?

- statistics increases linearly with  $E$  (cross section)  $\rightarrow$  increase rates (very important for anti-neutrinos)
- longer baseline  $\rightarrow$  enhance matter effects  $\rightarrow$  possibility to measure the sign of  $\Delta m^2_{13}$
- increase the energy  $\rightarrow$  easier to measure the spectral information in the oscillation signal  $\rightarrow$  important to reduce the intrinsic degeneracies
- Atmospheric background becomes negligible (this is a major background source in the low energy option)  $\rightarrow$  the bunching of the ions is not more a crucial issue

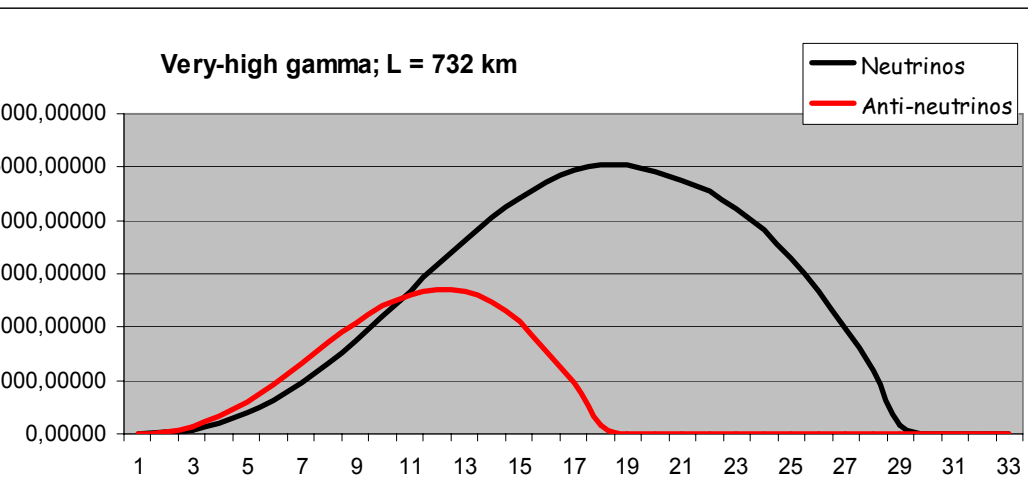


# Which $\gamma$ 's?

- Use a refurbished SPS with superconducting magnets to accelerate ions
  - Maximum  $\gamma \sim 600$
- Use the LHC to accelerate ions
  - Up to  $\gamma \sim 2488$  for  ${}^6\text{He}$  and 4158 for  ${}^{18}\text{Ne}$

Machine	Proton kinetic energy (GeV)	$\gamma(p)$	$\gamma({}^6\text{He}^{2+})$	$\gamma({}^{18}\text{Ne}^{10+})$
FNAL Booster	8	9.5	3.3	5.4
Main Injector	150	161	64	89
Tevatron	980	1045	349	581
BNL Booster	2	3.1	1.4	1.9
AGS	30	34	11	19
RHIC	250	268	89	149

In the US (see talk of S.Geer and APS meeting @ Snowmass, 28-30 Jun 04)



- Cross-sections are not an issue anymore ( $E_\nu > 1 \text{ GeV}$ )
- Megatonne water Cerenkov detectors are not needed anymore
- $\nu_e \rightarrow \nu_\tau$  becomes possible
- More challenging from the accelerator point of view

# EURISOL

## *Design Study*

- Total budget is 33293300 (9161900 from EU)
- Start date: 1 January 2005
- Objective: TDR for end of 2008
- Objective: TDR enabling the Nuclear physics and Neutrino physics communities to take a decision about a future facility
- 2009: Fix site and apply for EU construction project





# Beta-beam task

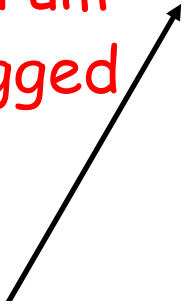
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- **Objective:** Study all components of a beta-beam facility above 100 MeV/u
- **Deliverable:** Conceptual Design Report (CDR) for a beta-beam facility
- **Participating institutes:** CERN, CEA, IN2P3, CLRC-RAL, GSI, MSL-Stockholm
- **Parameter group to define the conceptual design and follow the evolution of the beta-beam facility:** Higher intensities and higher gamma

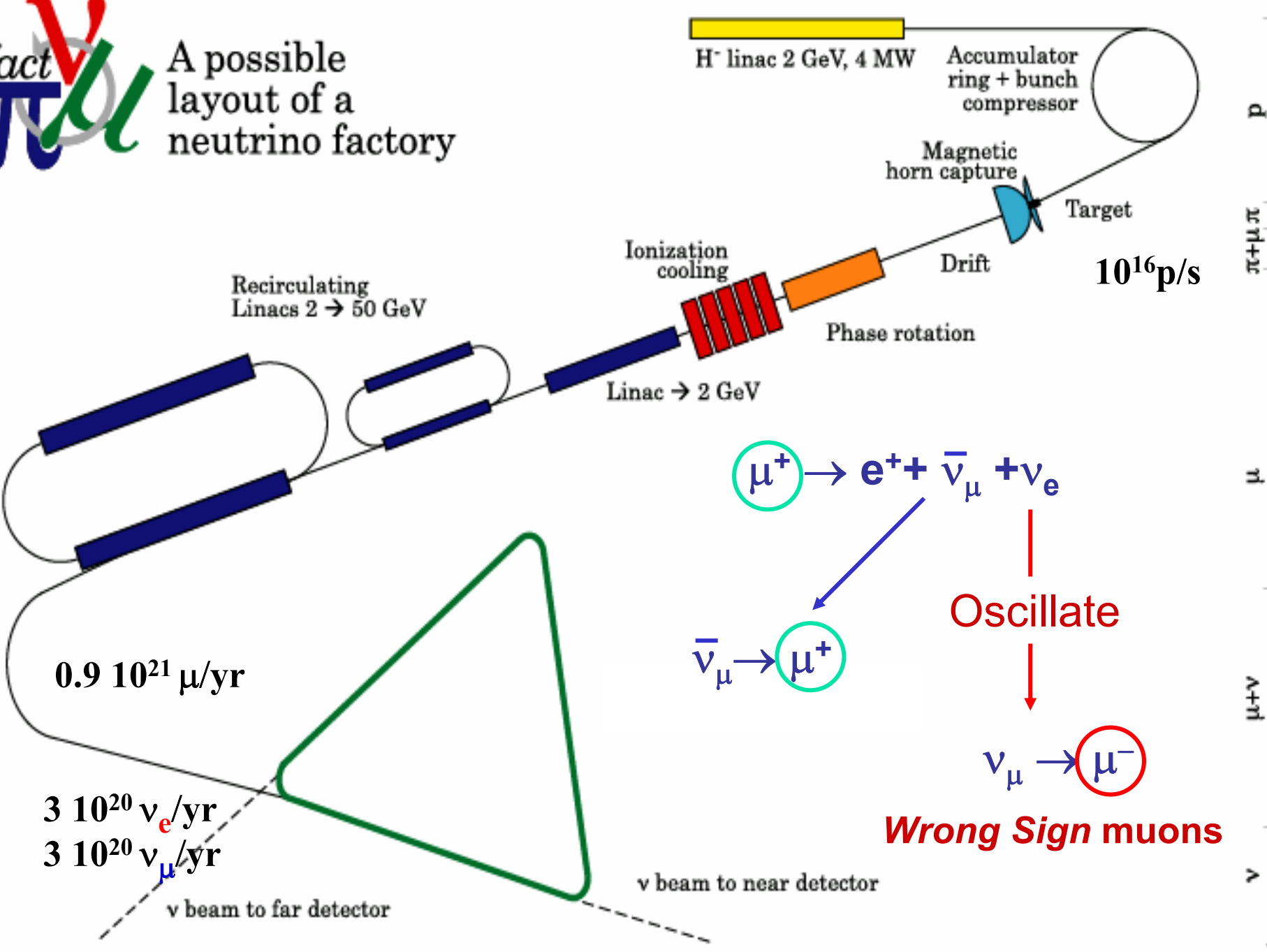


# Neutrino Factories

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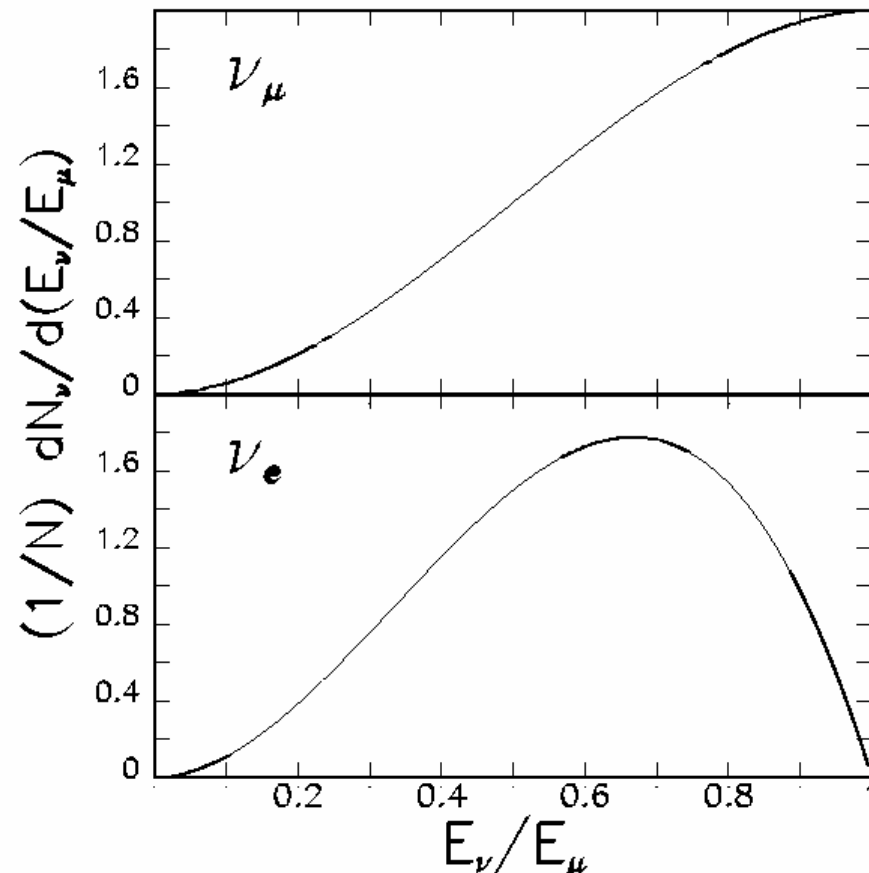
- The ultimate tool for probing neutrino oscillation, based on muon decays (NOT  $\pi$  DECAY !!!)
    - Enormous luminosity
    - Exceptional purity
    - Perfect knowledge of spectrum
    - Flavor of initial neutrino tagged by charge
  - Caveats:
    - Technical challenges to muon acceleration
    - Cost
- 
- Proton drivers
  - Targetry
  - Particle production measurements
  - RF manipulation
  - Cooling
  - Muon acceleration

# A possible layout of a neutrino factory



# Advantages of Muon Storage Ring

- Both  $\nu_e$  and  $\nu_\mu$  species in beam:
  - A way to get well understood, high-intensity source of  $\nu_e$ 's  
 $\nu_e \rightarrow \nu_\tau$  or  $\nu_e \rightarrow \nu_\mu$
- High intensity allows:
  - Probe small mixing angles
- Long distances
  - Start to see earth matter effects for oscillations involving  $\nu_e$ 's
  - Reach solar neutrino region with accelerator beams



# Comparison with Conventional $\nu$ Beam

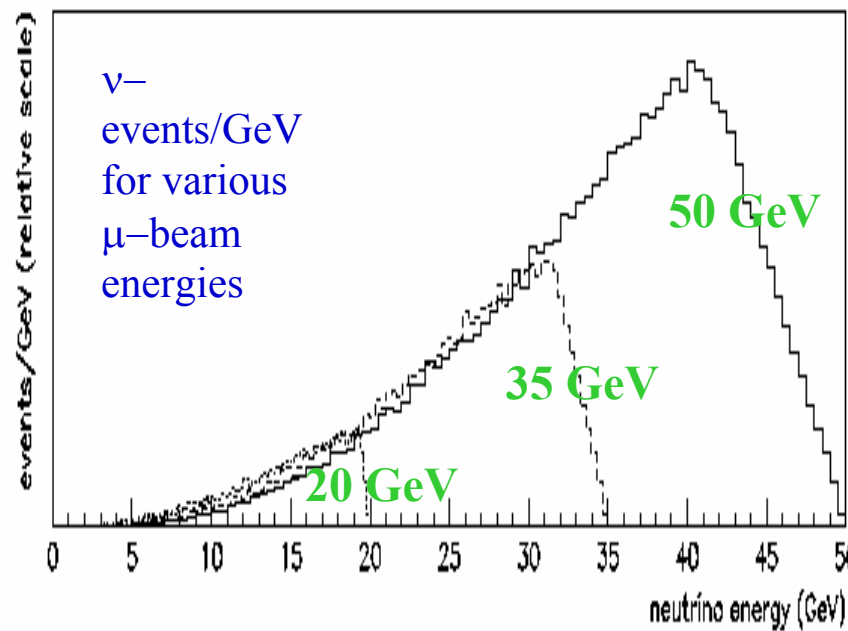
Experiment (L = 732 km)	CC Rates (per kt-yr)	
	$\nu_\mu$	$\bar{\nu}_e$
<b>MINOS (WBB)</b>		
Low Energy	458	1.3
Medium Energy	1439	0.9
High Energy	3207	0.9
<b><math>\mu</math>-ring</b>		
10 GeV	2200	1300
20 GeV	18 000	11 000
50 GeV	$2.9 \times 10^5$	$1.8 \times 10^5$
250 GeV	$3.6 \times 10^7$	$2.3 \times 10^7$

↓ Rate  
 $\sim E^3$

# $\nu$ - Factory Beam Parameters

## High Rate Beam:

- $10^{20} - 10^{21}$  muon decays/yr
- $\nu$  rates higher than conventional beams for  $E_{\text{storage}} > \sim 20 \text{ GeV}$
- Rate in detector  $\propto E^3 \Rightarrow$  High storage ring energy  $\sim 50 \text{ GeV}$





# Measurement Uncertainties at Neutrino Factory

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In general, MUCH SMALLER than  
Conventional Beams

Flux Uncertainties come from:

- current of muons in the ring
- divergence of muon beam
- Energy
- Polarization (usually unpolarized)

■ Cross Section Issues

- Energies are higher, in DIS regime
- Energy resolution good ( $>10\text{GeV}$ )
- $\nu$  production of charged particles well-modeled—less important

## At least 4 phases of Long Baseline experiments

2001

1) 2001-2010. K2K, Opera, Icarus, Minos.

Optimized to confirm the SuperK evidence of oscillation of atmospheric neutrinos through  $\nu_\mu$  disappearance or  $\nu_\tau$  appearance. They will have limited potential in measuring oscillation parameters. Not optimized for  $\nu_e$  appearance ( $\theta_{13}$  discovery).

2010

2) 2009-2015. T2K (approved), No $\nu$ a, Double Chooz. Optimized to measure  $\theta_{13}$  (Chooz  $\times$  20) through  $\nu_e$  appearance or  $\nu_e$  disappearance. Precision measure of the atmospheric parameters (1 % level). Tiny discovery potential for CP phase  $\delta$ , even combining their results.

2015

$10^{-1}$

$10^{-2}$

$10^{-3}$





# Parameters to Measure

**Atmospheric:**

$\Delta m_{23}^2$



$\theta_{23}$



**Solar:**

$\Delta m_{12}^2$



$\theta_{12}$



$\theta_{13}$



$\delta_{CP}$



**Mass  
hierarchy**





# Possible scenarios after first results of the planned experiments and implications

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- $\theta_{13}$  is so small ( $< 3^\circ$ ,  $\sin^2 2\theta_{13} \leq 0.01$ ) that all give null result
  - We need a “cheap” experiment to probe  $\sin^2 2\theta_{13}$  values down to  $O(0.001 - 0.0001)$
- $\theta_{13}$  is larger than  $3^\circ$  ( $\sin^2 2\theta_{13} \geq 0.01$ )
  - We need an experiment (or more than one) to
    - Measure  $\theta_{13}$  more precisely
    - Discover  $\delta$  (if not done yet) or precisely measure it
    - Measure the sign of  $\Delta m^2_{13}$
    - Measure  $\theta_{23}$  (is it  $\neq 45^\circ$ ?)
- NB Independently of the scenario the worsening of the experimental sensitivity due to the eightfold degeneracy has to be taken into account



# How to exploit high $\gamma$ BB?

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- Phase I exps give null result
  - See hep-ph/0405081 for a cheap and extremely sensitive to  $\theta_{13}$  experiment
- Phase I discover  $\theta_{13}$ 
  - See Nucl.Phys.B695:217-240,2004 for possible setups to search for  $\delta$
  - New ideas



# A proposal for a cheap experiment

Signal: an excess of horizontal muons in coincidence with the beam spill (possible thanks to the BB flavour composition)

✓ Number of unoscillated events:  
increase linearly with  $E$

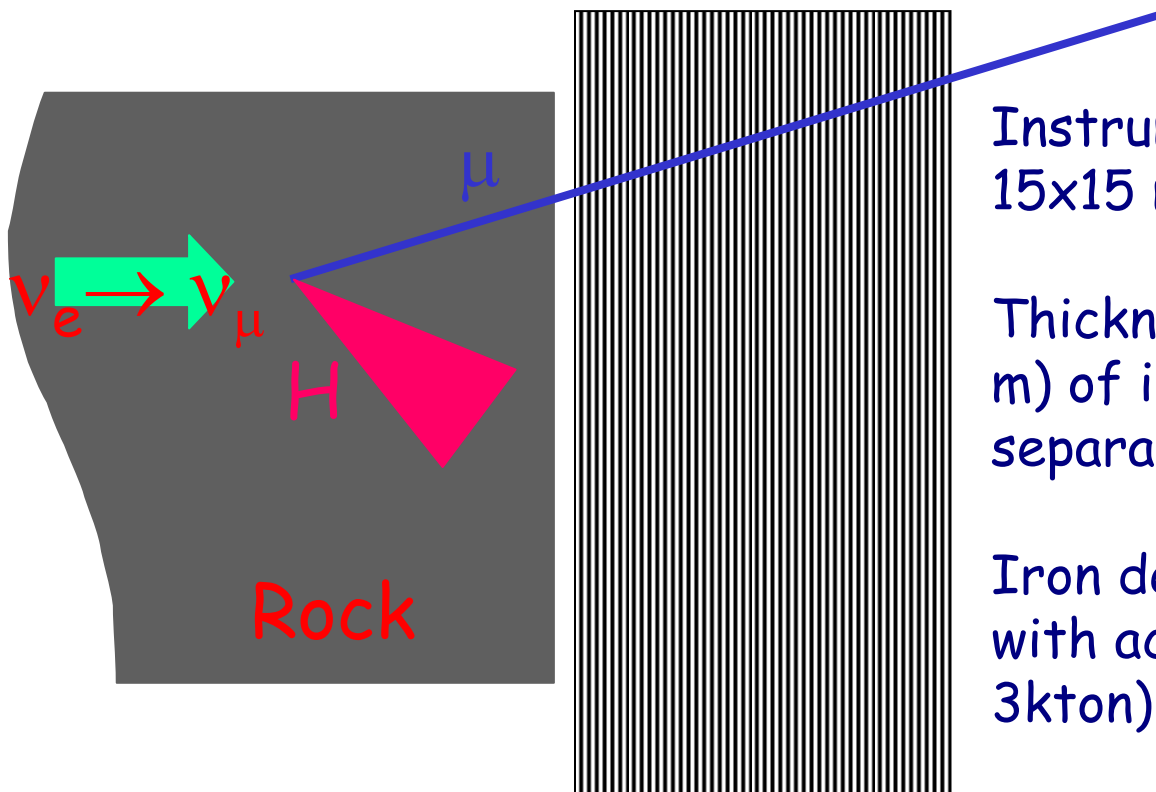
✓ Range of muons: increase linearly  
with  $E$  as well. The effective volume  
of rock contributing to the statistics  
increase linearly with  $E$

✓ The cost of the detector increase  
with the surface and not with the  
volume

We gain a quadratic increase of  
the sensitivity if we increase  $\gamma$   
and we reduce the detector cost  
by order of magnitudes!

We loose the possibility to fully  
reconstruct the events

# Schematic view of the detector



Instrumented surface:  
15x15 m<sup>2</sup> (one LNGS Hall)

Thickness: at least  $8\lambda_I$  (1.5 m) of iron for a good  $\pi/\mu$  separation

Iron detector interleaved with active trackers (about 3kton)



# A possible scenario: BB from CERN to Gran Sasso

- A cavern already exists at GS, but
    - Too small to host 40 kton WC or LAr detectors
    - On peak exp requires  $E_\nu \sim 1\text{-}2\text{ GeV}$  ( $\gamma = 350/580$ )  $\Rightarrow$  too small to efficiently exploit iron detectors
  - What happens if we consider  $\gamma > 1000$  (i.e. off-peak experiment)?
    - The oscillation probability decreases as  $\gamma^{-2}$
    - The flux increases as  $\gamma^2$
    - The cross-section and the effective rock volume increase both as  $\gamma$
    - Matter effects cancel out at leading order even if the baseline is large
- $\Rightarrow$  We recover the quadratic increase of sensitivity but we test now CP-even terms and no matter effects



# Event rate

## Beam assumptions

$1.1 \times 10^{18}$  decays per year of  $^{18}\text{Ne}$

$2.9 \times 10^{18}$  decays per year of  $^6\text{He}$

## Applied cuts

2 GeV energy cut in a  $20^\circ$  cone

100 % oscillated events/year:

$3 \times 10^4$  ( $\nu_e$  @  $\gamma=2500$ )

$0 \times 10^4$  (anti  $\nu_e$  @  $\gamma=1500$ )

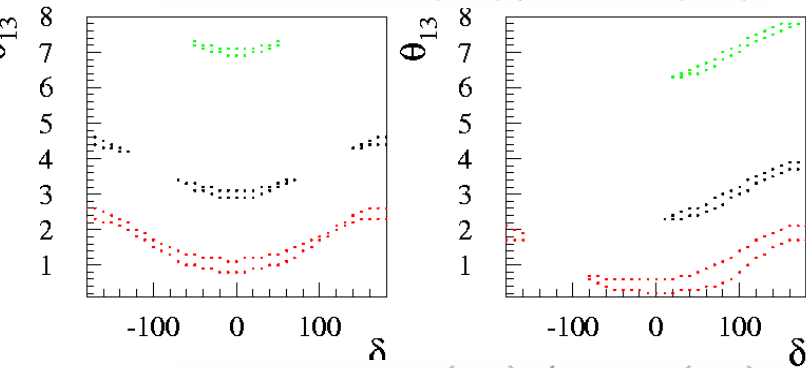
$9 \times 10^5$  ( $\nu_e$  @  $\gamma=4158$ )

$1 \times 10^5$  (anti  $\nu_e$  @  $\gamma=2488$ )

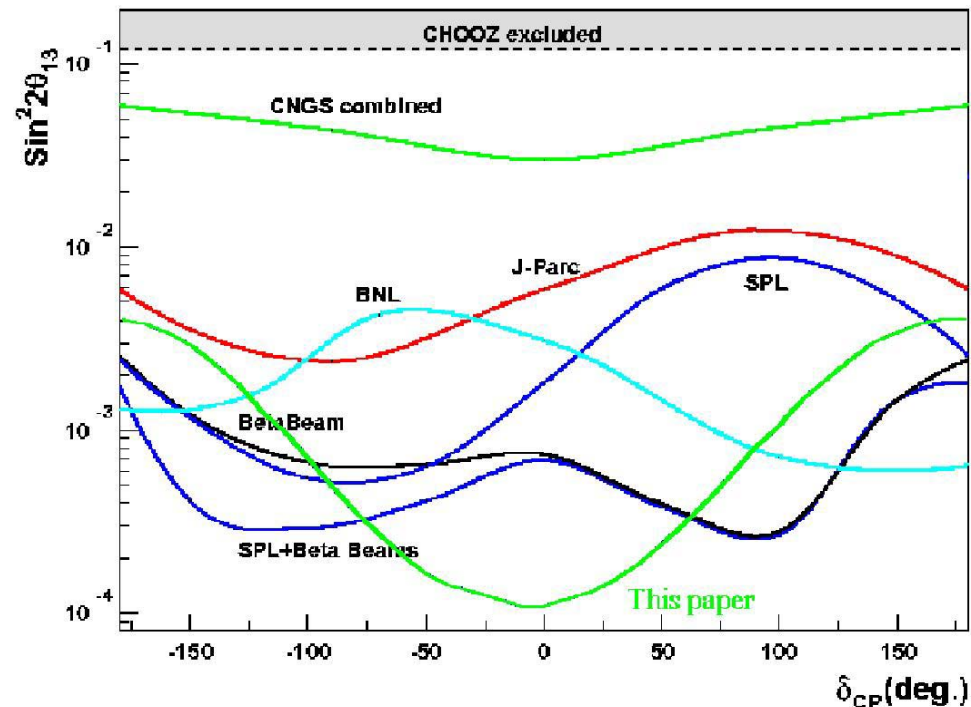
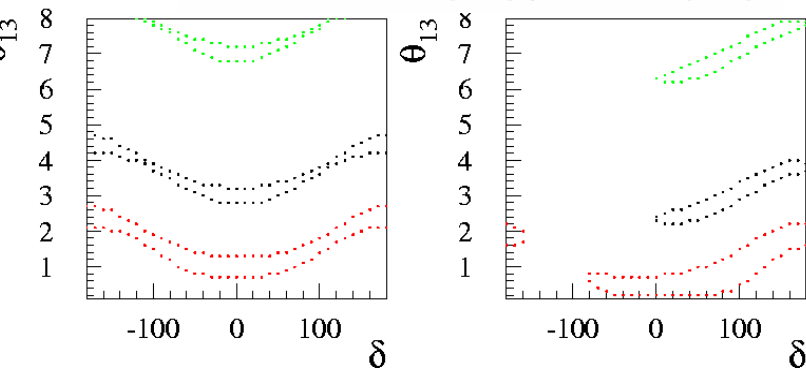
Detector	$\gamma$	$\nu_e \rightarrow \nu_\mu$	$\pi$	$\mu$	charm
$B = 0$ T	2500 ( $\nu_e$ )	1.5	0.5	11.6	20.2
$B = 0$ T	1500 ( $\bar{\nu}_e$ )	0.8	0.02	3.5	1.5
$B = 0$ T	4158 ( $\nu_e$ )	4.9	4.6	153.4	357.1
$B = 0$ T	2488 ( $\bar{\nu}_e$ )	3.2	0.3	15.4	37.1
$B \sim 1$ T	2500 ( $\nu_e$ )	1.5	0.2	5.8	0.2
$B \sim 1$ T	1500 ( $\bar{\nu}_e$ )	0.8	0.01	1.8	0.01
$B \sim 1$ T	4158 ( $\nu_e$ )	4.9	1.8	64.8	3.6
$B \sim 1$ T	2488 ( $\bar{\nu}_e$ )	3.2	0.1	7.8	0.4

# Sensitivity of a "massless" detector located 730 km from a (very)high- $\gamma$ BB

$$\gamma = 4158(\nu_e)/2488(\bar{\nu}_e)$$



$$\gamma = 2500(\nu_e)/1500(\bar{\nu}_e)$$

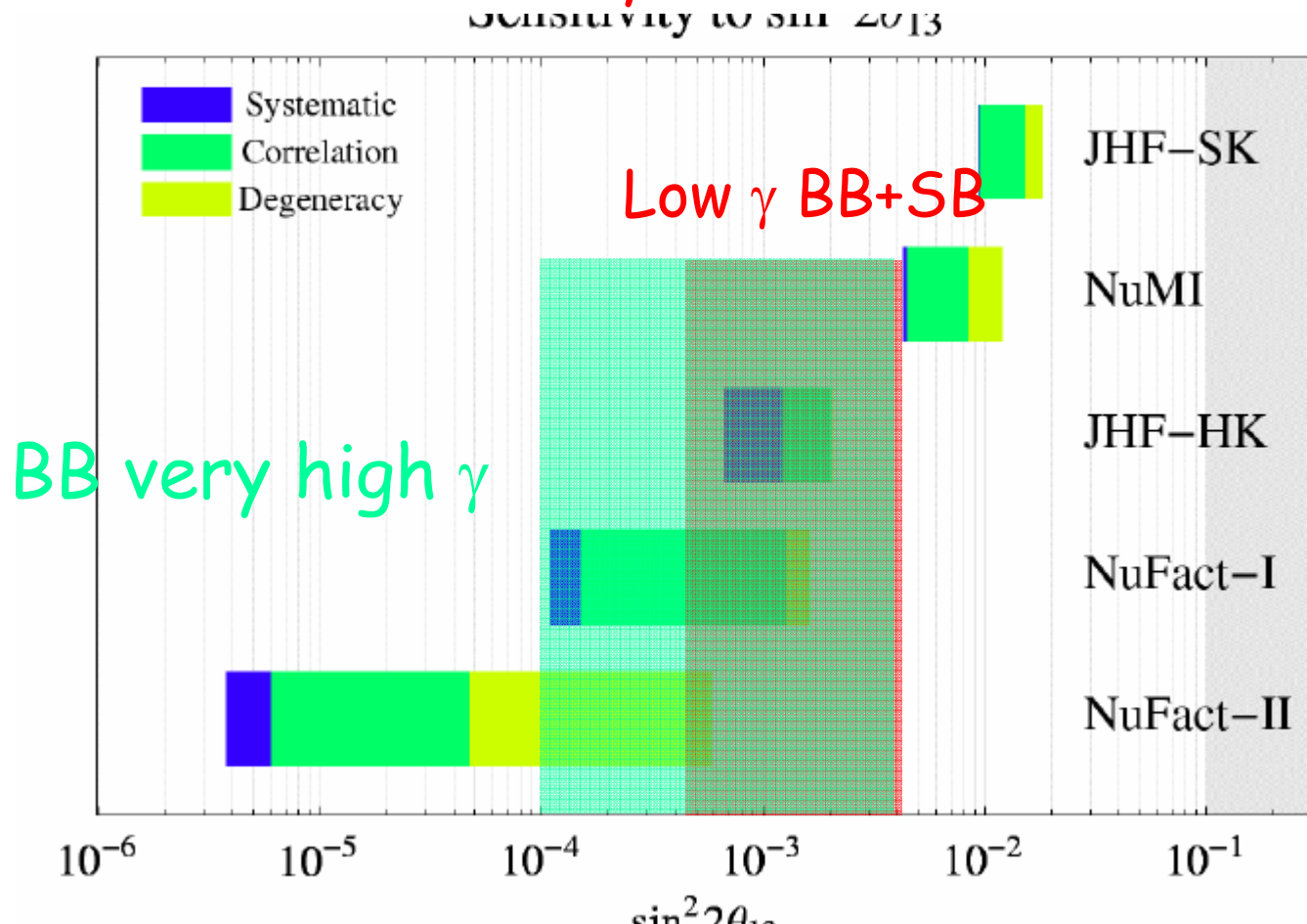


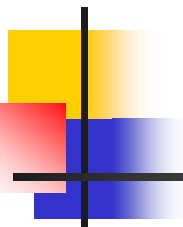
test  $\sin^2 2\theta_{13}$  values down to  $10^{-3}$ - $10^{-4}$ !!!



# Comparison of very-high $\gamma$ BB with some of future projects

In case of null result very difficult to build new facilities!





## Two setups studied for the medium/high $\gamma$ options

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- Medium (350) and high (1500)  $\gamma$  for medium (730 km) and far (3000 km) baselines
- Water detector (UNO) like; 1 Mton mass. Includes full simulation of efficiencies and backgrounds (only statistical study for high gamma option)
- Running time 10 years
- Full analysis (including the eightfold degeneracy, all systematics on cross-sections, detector, beam, performance at small  $\theta_{13}$ , etc.) still to be done

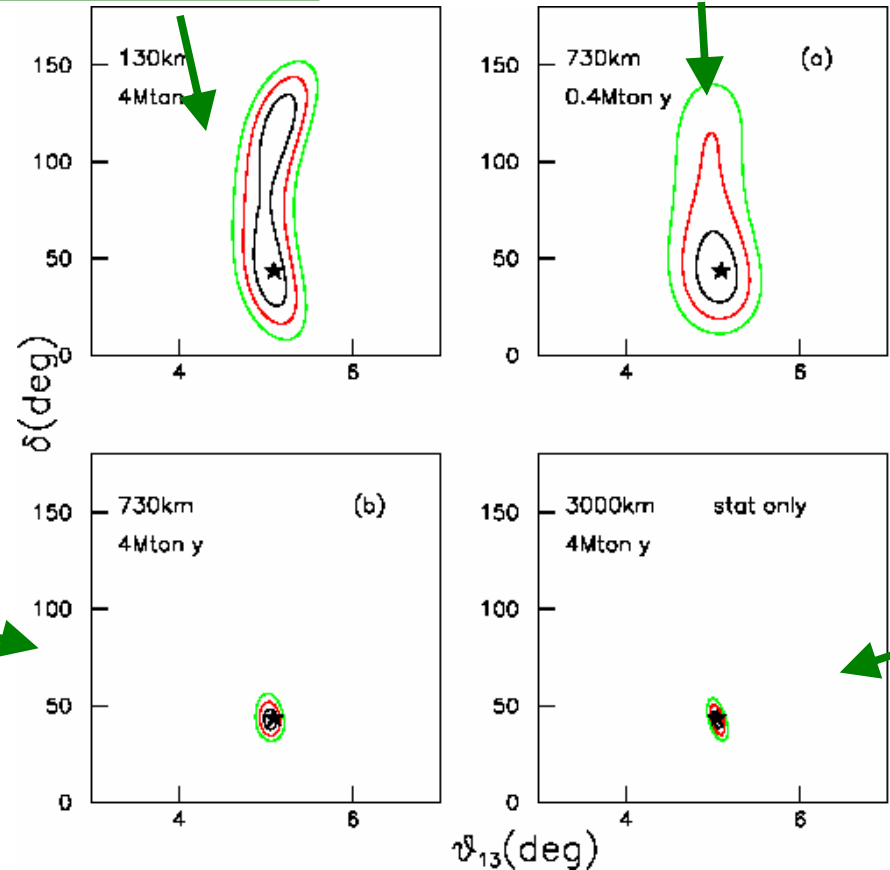
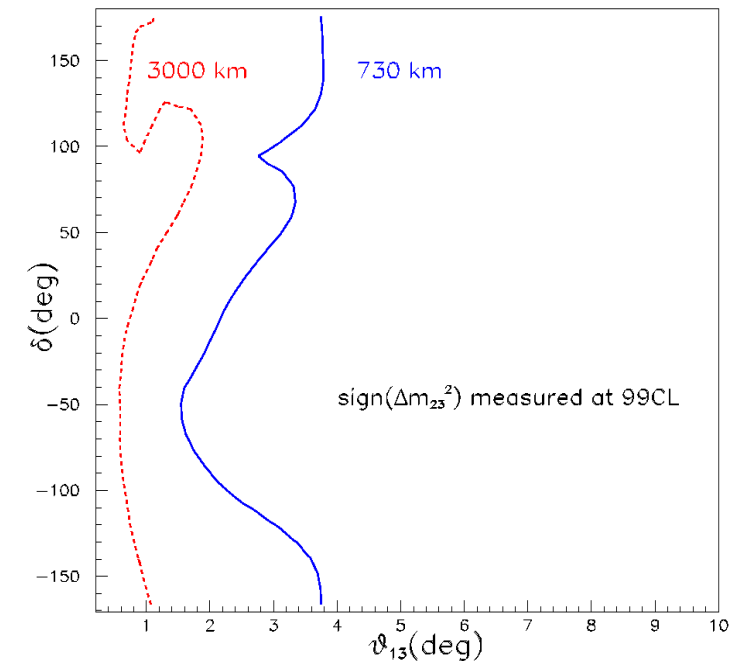
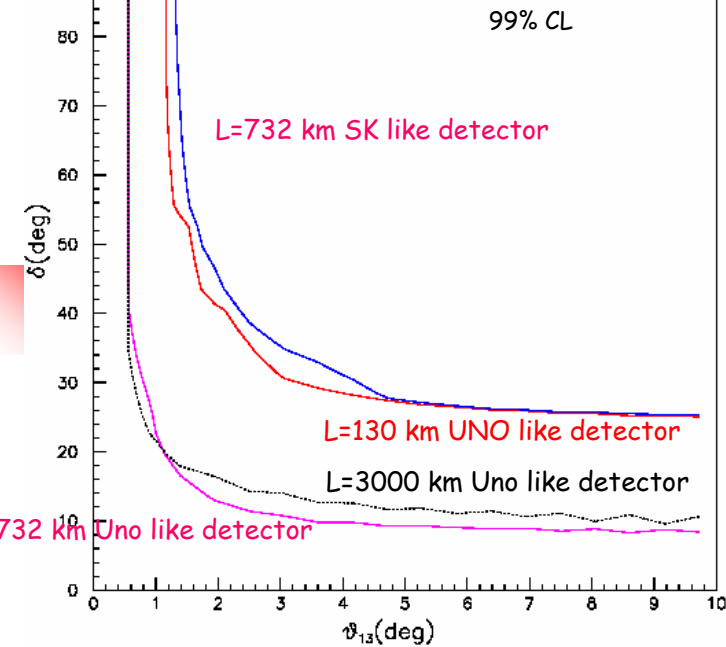
# Results

4 Mton/y WC detector @ 3000 km  
 $\gamma=1550$  ( $^6\text{He}$ ) / 2500 ( $^{18}\text{Ne}$ )

4 Mton/y WC @ 730 km  
 $\gamma=350$  ( $^6\text{He}$ ) / 580 ( $^{18}\text{Ne}$ )

40 Kton/y WC @ 730 km  
 $\gamma=350$  ( $^6\text{He}$ ) / 580 ( $^{18}\text{Ne}$ )

Baseline option  
 (Frejus)





# Comments

The idea of medium/(very) high- $\gamma$  BB is very appealing

Whatever  $\gamma$  (medium, high, very-high) we consider its performance is better than the low one

The medium scenario has been put forward in Nucl.Phys.B695:217-240,2004 to measure  $\theta_{13}$ ,  $\delta$  and the sign of  $\Delta m^2_{13}$ , but more studies are needed to fully exploit its potential (i.e. the  $\theta_{23}$  ambiguity)

However, we think this is not the optimal solution

- It foresees the construction of a 1 Mton detector!
- There are no place in the world able to host it
- It is very expensive, so to risky to build if phase I exps give null results

The optimal solution is the very-high  $\gamma$  scenario

- In case of null result of phase I exps it allows a cheap investigation of very small values of  $\sin^2 2\theta_{13}$  (see hep-ph/0405081)
- In case of positive result of phase I exps it allows a complete study of the PMNS matrix through different channels, see next slides for details
- On top of that it makes possible the usage of magnetized calorimeters which are smaller (40 kton  $\rightarrow$  about  $10^4$  m<sup>3</sup>) than WC detectors (1 Mton  $\rightarrow$  about  $10^6$  m<sup>3</sup>)  $\rightarrow$  cheaper (easier) civil engineer costs

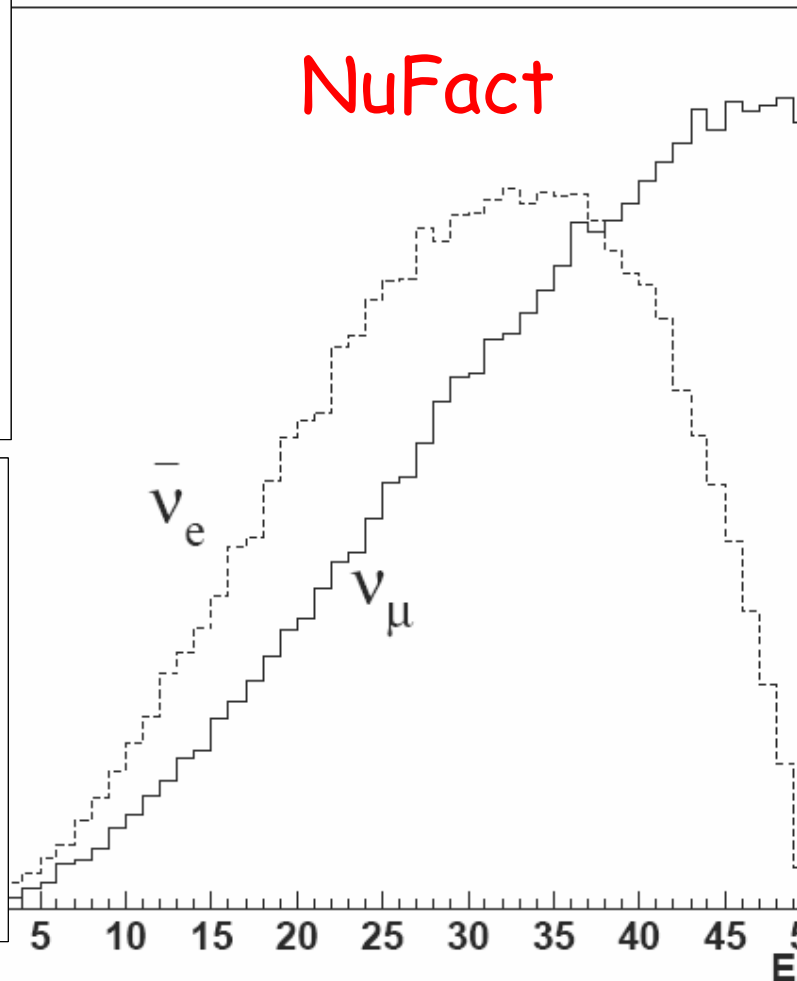
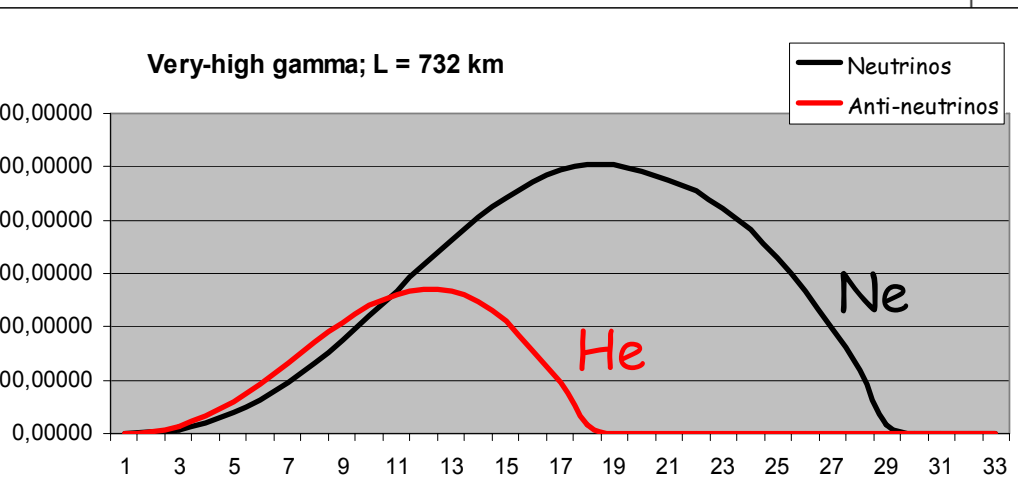
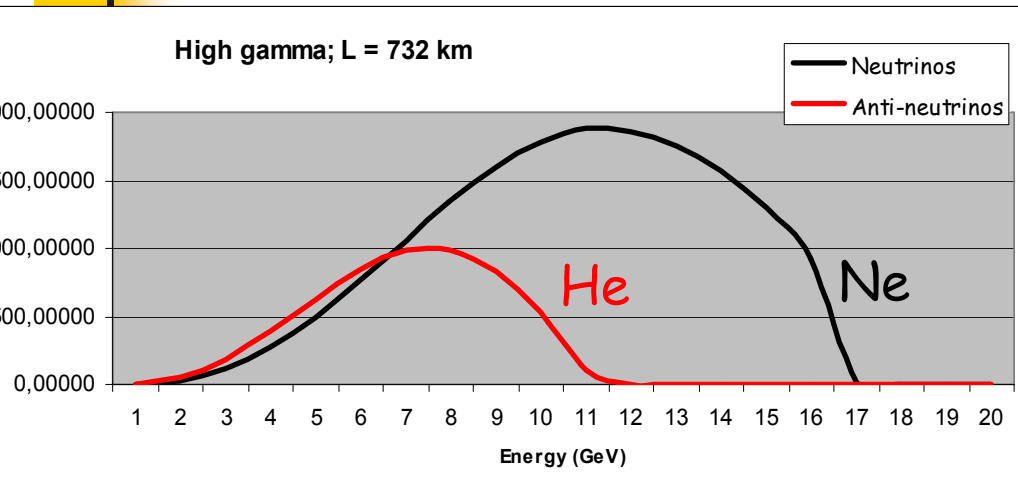


# Preliminary studies/ideas on how to use the very-high $\gamma$ BB

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A. Donini, PM, S. Rigolin, ...

# BB vs NuFact spectra



# Expected rates (1ktonx1year)

	NuFact		BB	
	$\nu_e$	anti- $\nu_\mu$	$\nu_e$	anti- $\nu_e$
L=730 km	$296 \times 10^3$	$176 \times 10^3$	---	---
L=730 km $\gamma=2500/1500$	---	---	$17 \times 10^3$	$5.5 \times 10^3$
L=730 km $\gamma=4158/2488$	---	---	$77 \times 10^3$	$25 \times 10^3$
L=3000 km	$18 \times 10^3$	$11 \times 10^3$	---	---
L=3000 km $\gamma=2500/1500$	---	---	$1.0 \times 10^3$	$0.3 \times 10^3$
L=3000 km $\gamma=4158/2488$	---	---	$4.6 \times 10^3$	$1.5 \times 10^3$

**NB** There is less than a factor 10 difference in the #evts  
 BB allows simultaneous run with  $\nu$  and anti- $\nu$ , while NuFact does not

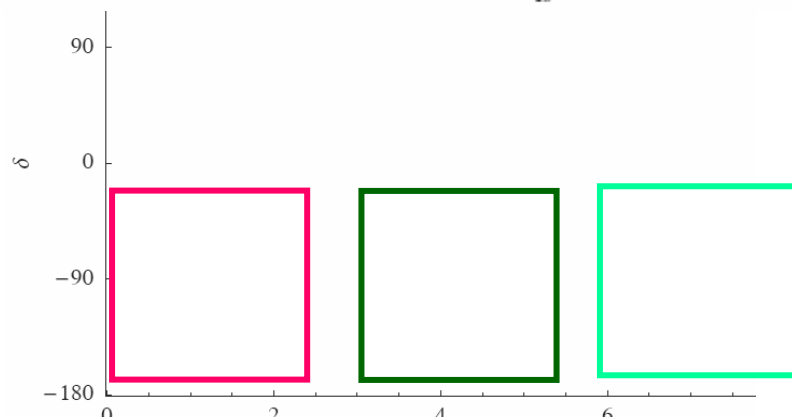
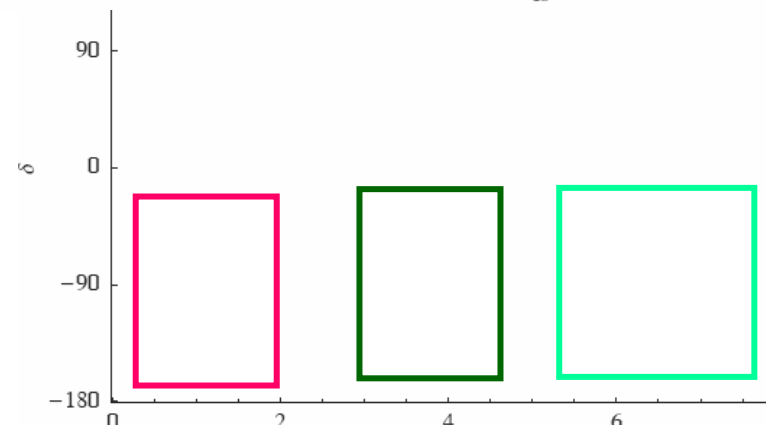
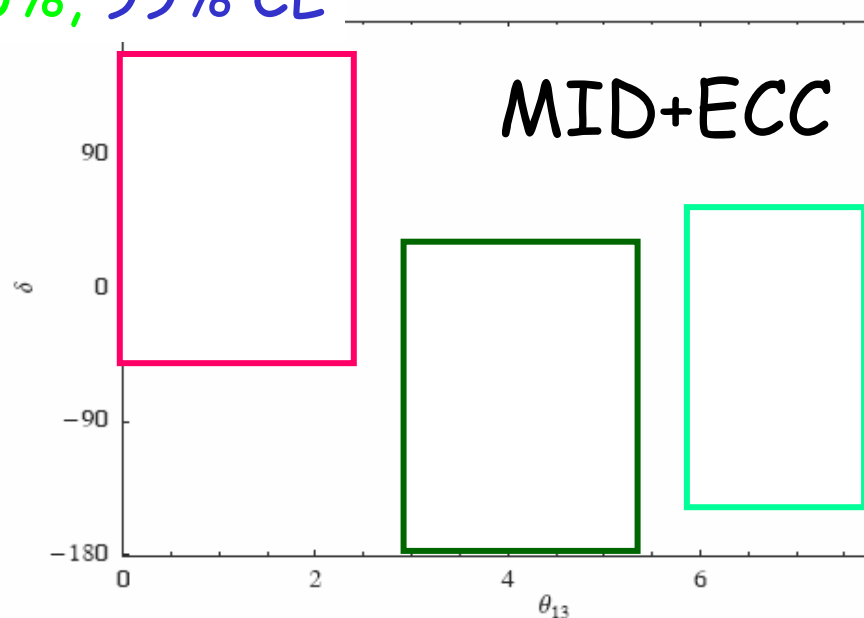
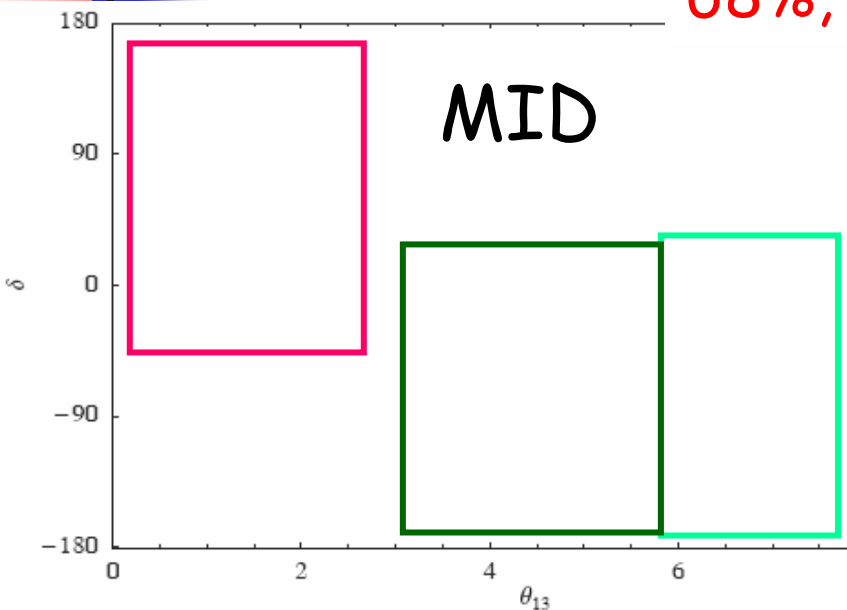
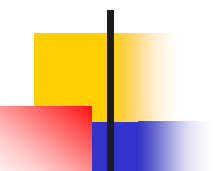


# Potentiality of a very-high $\gamma$ BB

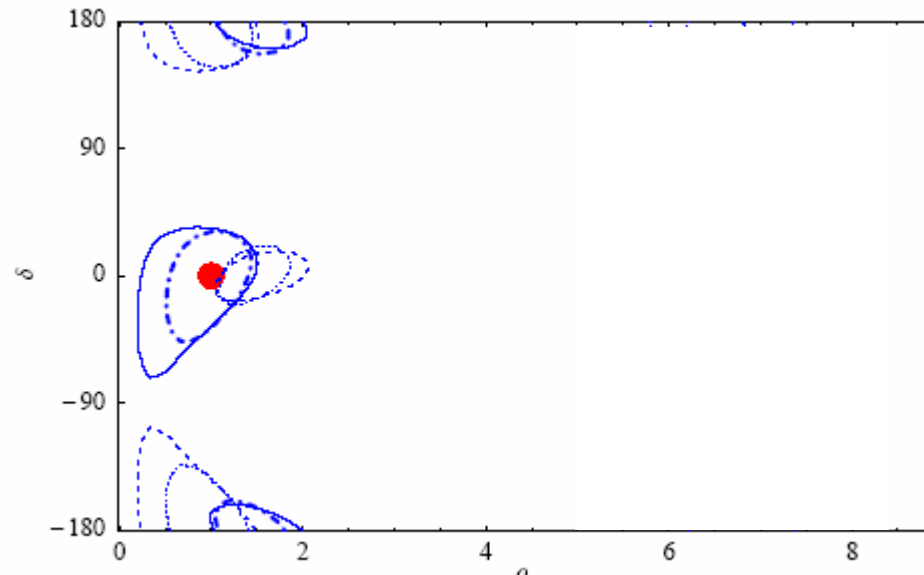
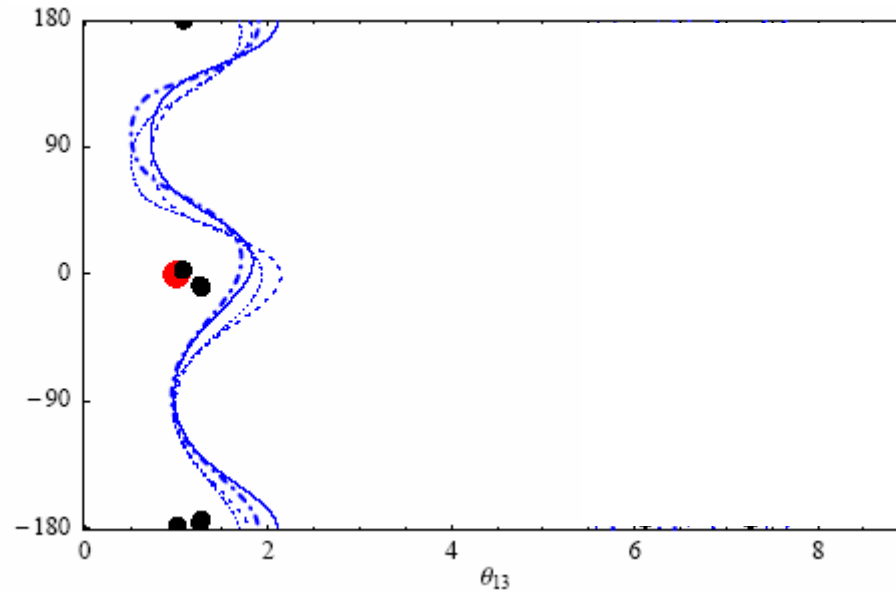
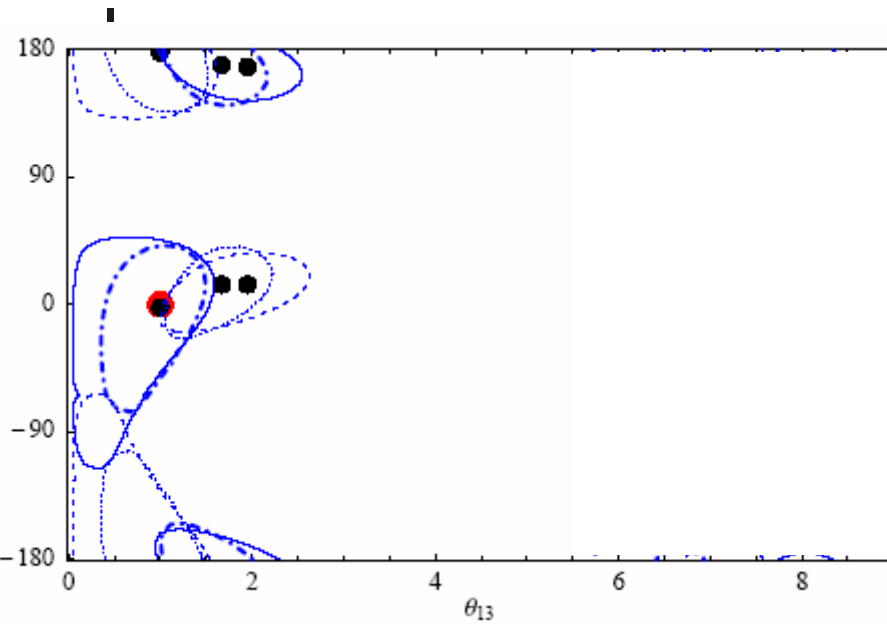
- Simultaneous search for  $\nu_e \rightarrow \nu_\mu$  (golden) and  $\nu_e \rightarrow \nu_\tau$  (silver) channels
  - This combination is highly efficient in removing the intrinsic and the sign degeneracy (see A.Donini, D.Meloni, P.Migliozzi Nucl.Phys.B646:321-349,2002)
- Simultaneous search for  $\nu$  and anti- $\nu$  channels (i.e. 1 year BB  $\equiv$  2 years NuFact)
- Detectors: 40kton magnetized iron detector (MID) at 3000km;  $\geq 5$ kton ECC detector at 730km
- The physics potential of this setup is currently under study as well as its comparison with a NuFact



Very preliminary results  
at a high- $\gamma$  BB  
with golden blue silver channels  
68%, 90%, 99% CL



# Parameter extraction in presence of signal (11) with a low $\gamma$ BB plus a SB



Continuous line: intrinsic degeneracy

Dashed line: sign ambiguity

Dot-dashed line: octant ambiguity

Dotted line: mixed ambiguity



Solving all degeneracy's needs  
"gold, silver and water"

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# The Detectors Setup

50 GeV Neutrino Factory located at CERN

Three detectors of different design:

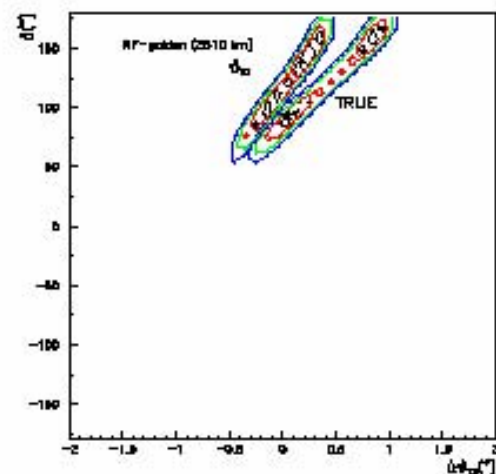
- 40 Kton Magnetized iron detector (MID)  
 $L = 2810$  Km (Canary Islands)  
A. Cervera *et al.*,  
Nucl. Instr. Meth. A 451 (2000) 123; NuFact99, Lyon
- 400 Kton Water Cherenkov (WC)  
 $L = 130$  Km (Frejus)  
A. Blondel *et al.*,  
Nucl. Instr. Meth. A 503 (2001) 173; NuFact01, Tsukuba
- 4 Kton Emulsion Cloud Chamber (ECC)  
 $L = 732$  Km (Gran Sasso) or  $L = 2810$  Km  
D. Autiero *et al.*, hep-ph/0305185; NuFact03, New York

# One detector

Consider the NuFact golden channel:  
best option for one detector, with baseline  $L = 2810$   
(no sign degeneracies for  $\theta_{13} \geq 1^\circ$ ).

A. Cervera *et al.*, hep-ph/0002108

- 40 Kton MID



$$\bar{\theta}_{13} = 2^\circ \quad \bar{\delta} = 90^\circ$$

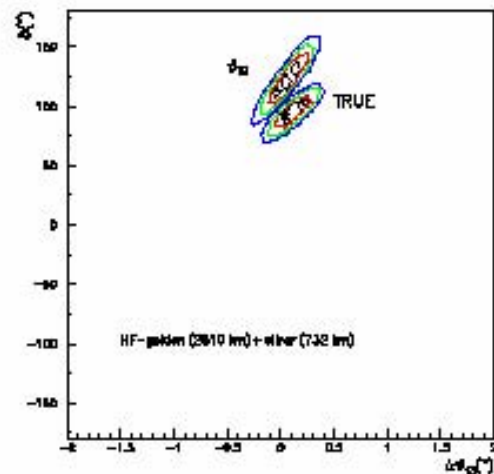
# Two detectors

You can now add a second detector.

We can take advantage of the NuFact silver channel...

A. Donini *et al.*, hep-ph/0206034

- 40 Kton MID
- 4 Kton ECC



$$\bar{\theta}_{13} = 2^\circ \quad \bar{\delta} = 90^\circ$$

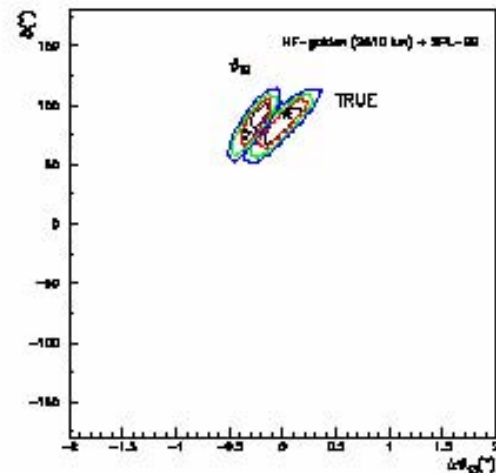


# Two detectors

... or of the Superbeam-driven water Cherenkov.

J. Burguet-Castell *et al.*, hep-ph/0207080

- 40 Kton MID
- 400 Kton WC



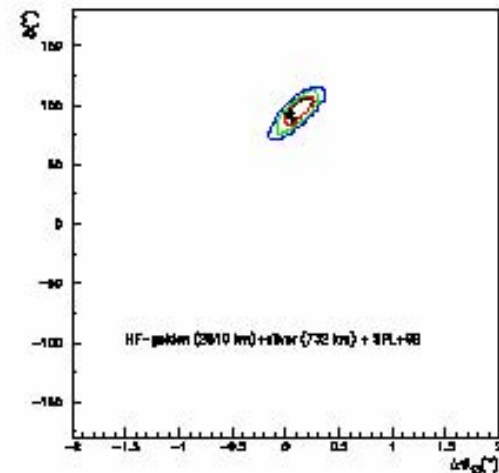
$$\bar{\theta}_{13} = 2^\circ \quad \bar{\delta} = 90^\circ$$

# The Three Detectors

However, the very best possibility is  
to combine the three detectors in their **FULL GLORY**.

.... Halleluja....

- 40 Kton MID
- 4 Kton ECC
- 400 Kton WC



$$\bar{\theta}_{13} = 2^\circ \quad \bar{\delta} = 90^\circ$$



# Sensitivity limits

Statement:

All degeneracies solved for  $\theta_{13} \geq 1^\circ$   
( $\sin^2 2\theta_{13} \geq 0.001$ )

However, three comments are in order:

- 4 Kton ECC: statistical limit at  $\sim 1^\circ$   
(P. Migliozzi's talk and hep-ph/0305185)
- 40 Kton MID: effect of systematics?  
(P. Migliozzi's talk and hep-ph/0305185)
- 400 Kton WC: can we reduce its mass?  
how this affects our results?

A conservative statement:

All degeneracies solved for  $\theta_{13} > 1^\circ$